# Pseudo-symmetric spaces of constant type in dimension three - elliptic spaces 

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Riassunto: Diamo una classificazione quasiesplicita delle varietà riemanniane 3dimensionali cosiddette "ellittiche", cioè con autovalori di Ricci $\rho_{1}=\rho_{2} \neq \rho_{3}, \rho_{3}=$ costante $\neq 0$. (Nel caso non ellittico il problema é stato risolto esplicitamente in [14]). Le classi locali di isometria delle metriche di tali varietà dipendono essenzialmente da tre funzioni arbitrarie di due variabili. Diamo anche un esempio di famiglie esplicite di metriche che dipendono da due funzioni arbitrarie di due variabili.

Abstract: We give a quasiexplicit classification of three-dimensional Riemannian manifolds with Ricci eigenvalues $\rho_{1}=\rho_{2} \neq \rho_{3}, \rho_{3}=$ constant $\neq 0$; which are called "elliptic". (In the nonelliptic cases the problem was solved explicitly in [14]). The local isometry classes of metrics of such manifolds depend on essentially three arbitrary functions of two variables. We also give an example of an explicit family of metrics depending on two arbitrary functions of two variables.

## - Introduction

According to [5], a Riemannian manifold $(M, g)$ is said to be pseudosymmetric if the following formula holds for arbitrary vector fields $X$ and

[^0]$Y$ on $M$ :
\[

$$
\begin{equation*}
R(X, Y) \cdot R=F((X \wedge Y) \cdot R) \tag{0.1}
\end{equation*}
$$

\]

where
a) $R$ denotes the Riemannian curvature tensor of type $(1,3)$ on $(M, g)$ and

$$
R(X, Y)=\left[\nabla_{X}, \nabla_{Y}\right]-\nabla_{[X, Y]}
$$

denote the corresponding curvature transformations,
b) $X \wedge Y$ denotes the endomorphism of the tangent bundle $T M$ defined by

$$
\begin{equation*}
(X \wedge Y) Z=g(Y, Z) X-g(X, Z) Y \tag{0.2}
\end{equation*}
$$

c) $F$ is a smooth function on $M$,
d) the dot in each side of the formula (0.1) denotes the derivation on the tensor algebra of $T M$ induced by an endomorphism of this tangent bundle.
We call a pseudo-symmetric space $(M, g)$ of constant type if $F=$ $\tilde{c}=$ constant. According to [4] we have the following characterization in dimension three (see [14] for more details):

Proposition 0.1. A three-dimensional Riemannian manifold ( $M, g$ ) is pseudo-symmetric of constant type $F=\tilde{c}$ if and only if its principal Ricci curvatures $\rho_{1}, \rho_{2}$ and $\rho_{3}$ locally satisfy the following conditions (up to a numeration):
(i) $\rho_{3}=2 \tilde{c}$,
(ii) $\rho_{1}=\rho_{2}$ everywhere.

We are not interested in the case when $(M, g)$ is a space of constant curvature and therefore we assume always $\rho_{1}=\rho_{2} \neq \rho_{3}$.

If $\tilde{c}=0$, and hence $F=0$, we obtain a definition of semi-symmetric space. The theory of semi-symmetric spaces has been developed in [17], [18], [19], [9], [6], [1], [2] and especially in the book [3]. For the threedimensional case, see the explicit classification in [9], [6] and [3, Chapter 6].

For $\tilde{c} \neq 0$, the present authors made an explicit classification in [14] for the so-called "asymptotically foliated" (or "non-elliptic") spaces in dimension three. (See section 4 for the terminology.) The aim of this paper is to treat the more complicated "elliptic" spaces in the full generality.

Let us mention that the first author in [10] solved the special case when $\rho_{1}=\rho_{2}$ is a constant ${ }^{(*)}$, and the present authors treated in [13] a more general case-here $\rho_{1}=\rho_{2}$ is supposed to be constant along each trajectory of the principal Ricci curvature $\rho_{3}$. The basic methods of [9], [10] and [13] are used also here but the corresponding calculations became more complicated. A computer check (the software "Mathematica" by Wolfram Research Inc.) was also used during this work.

## 1 - The basic system of partial differential equations for the problem

Let $(M, g)$ be a three-dimensional Riemannian manifold whose Ricci tensor $\hat{R}$ has eigenvalues $\rho_{1}=\rho_{2} \neq \rho_{3}$ with nonzero constant $\rho_{3}$. Choose a neighborhood $\tilde{U}$ of a fixed point $m \in M$ and a smooth vector field $E_{3}$ of unit eigenvectors corresponding to the Ricci eigenvalue $\rho_{3}$ in $\tilde{U}$. Let $S: D^{2} \rightarrow \tilde{U}$ be a surface through $m$ which is transversal with respect to all trajectories generated by $E_{3}$ at all cross-points and not orthogonal to such a trajectory at $m$. (The vector field $E_{3}$ determines an orientation of $S$.) Then there is a normal neighborhood $U$ of $m, U \subset \tilde{U}$, with the property that each point $p \in U$ is projected to exactly one point $\pi(p) \in S$ via some trajectory. We fix any local coordinate system $(w, x)$ on $S$ and then a local coordinate system $(w, x, y)$ on $U$ such that the values $w(p)$ and $x(p)$ are defined as $w(\pi(p))$ and $x(\pi(p))$, respectively, for each point $p \in U, y(p)$ is the oriented length $d^{+}(\pi(p), p)$ of the trajectory joining $p$ with $\pi(p)$. Then $E_{3}=\partial / \partial y$ can be extended in $U$ to an orthonormal moving frame $\left\{E_{1}, E_{2}, E_{3}\right\}$. Let $\left\{\omega^{1}, \omega^{2}, \omega^{3}\right\}$ be the corresponding dual coframe. Then the $\omega^{i}$ are of the form

$$
\begin{equation*}
\omega^{i}=a^{i} d w+b^{i} d x, \quad i=1,2, \quad \omega^{3}=d y+H d w+G d x \tag{1.1}
\end{equation*}
$$

The Ricci tensor $\hat{R}$ expressed with respect to $\left\{E_{1}, E_{2}, E_{3}\right\}$ has the form $\hat{R}_{i j}=\rho_{i} \delta_{i j}$. Because each $\rho_{i}$ is expressed through the sectional curvature
(*) See [11], [12], [15], [16] and [3], Chapter 12, for the related topics.
$K_{i j}$ by the formula $\rho_{i}=\hat{R}_{i i}=\sum_{j \neq i} K_{i j}$, there exist a function $k=$ $k(w, x, y)$ of the variables $w, x$ and $y$, and a constant $\tilde{c}$ such that

$$
\begin{equation*}
K_{12}=k, \quad K_{13}=K_{23}=\tilde{c}, \quad \rho_{1}=\rho_{2}=k+\tilde{c}, \quad \rho_{3}=2 \tilde{c} \tag{1.2}
\end{equation*}
$$

Define now the components $\omega_{j}^{i}$ of the connection form by the standard formulas

$$
\begin{equation*}
d \omega^{i}-\sum_{j} \omega^{j} \wedge \omega_{j}^{i}=0, \quad \omega_{j}^{i}+\omega_{i}^{j}=0, \quad i, j=1,2,3 \tag{1.3}
\end{equation*}
$$

Because the Riemannian curvature tensor satisfies $R_{i j k l}=0$ whenever at least three of the indices $i, j, k$ and $l$ are distinct, the formulas (1.2) are equivalent to

$$
\left\{\begin{array}{l}
d \omega_{2}^{1}+\omega_{3}^{1} \wedge \omega_{2}^{3}=k \omega^{1} \wedge \omega^{2}  \tag{1.4}\\
d \omega_{3}^{1}+\omega_{2}^{1} \wedge \omega_{3}^{2}=\tilde{c} \omega^{1} \wedge \omega^{3} \\
d \omega_{3}^{2}+\omega_{1}^{2} \wedge \omega_{3}^{1}=\tilde{c} \omega^{2} \wedge \omega^{3}
\end{array}\right.
$$

Next, differentiate the equations (1.4) and substitute from (1.4). We obtain easily

$$
\begin{equation*}
\omega_{3}^{1} \wedge \omega^{1} \wedge \omega^{2}=0, \quad \omega_{3}^{2} \wedge \omega^{1} \wedge \omega^{2}=0 \tag{1.5}
\end{equation*}
$$

and

$$
\begin{equation*}
d\left((k-\tilde{c}) \omega^{1} \wedge \omega^{2}\right)=0 \tag{1.6}
\end{equation*}
$$

The relations (1.5) mean that $\omega_{3}^{1}$ and $\omega_{3}^{2}$ are linear combinations of $\omega^{1}$ and $\omega^{2}$ only, and from the third equation of (1.3) it follows that $d \omega^{3}$ is a multiple of $\omega^{1} \wedge \omega^{2}$, i.e., a multiple of $d w \wedge d x$. Then (1.1) implies that the functions $G$ and $H$ are independent of $y$.

Now, there is a local coordinate system $(\bar{w}, \bar{x}, y)$ (possibly in a smaller neighborhood of $m$ ) such that $\bar{w}=\bar{w}(w, x)$ and $\bar{x}=\bar{x}(w, x)$ are functions of $w$ and $x$, and

$$
\begin{equation*}
\omega^{1}=P^{1} d \bar{w}+Q^{1} d \bar{x}, \quad \omega^{2}=P^{2} d \bar{w}+Q^{2} d \bar{x}, \quad \omega^{3}=d y+\bar{H}(\bar{w}, \bar{x}) d \bar{w} \tag{1.7}
\end{equation*}
$$

Indeed, because the surface $S$ is not orthogonal to the vector field $E_{3}$ at $m$, the Pfaffian form $H d w+G d x$ from (1.1) is nonzero in a neighborhood of $m$ in $M$. Then we define $\bar{w}=\bar{w}(w, x)$ as a potential function of the Pfaffian equation $H d w+G d x=0$, and the second function $\bar{x}=\bar{x}(w, x)$ can be defined as an arbitrary smooth function which is functionally independent of $\bar{w}$. In addition, there are new Pfaffian forms $\tilde{\omega}^{1}$ and $\tilde{\omega}^{2}$ such that $\left(\tilde{\omega}^{1}\right)^{2}+\left(\tilde{\omega}^{2}\right)^{2}=\left(\omega^{1}\right)^{2}+\left(\omega^{2}\right)^{2}$ and $\tilde{\omega}^{1}$ does not involve the differential $d \bar{x}$. We can summarize:

Proposition 1.1. In a normal neighborhood of any point $m \in M$ there exist an orthonormal coframe $\left\{\omega^{1}, \omega^{2}, \omega^{3}\right\}$ and a local coordinate system $(w, x, y)$ such that

$$
\begin{equation*}
\omega^{1}=f d w, \quad \omega^{2}=A d x+C d w, \quad \omega^{3}=d y+H d w \tag{1.8}
\end{equation*}
$$

Here $f, A$ and $C$ are smooth functions of the variables $w, x$ and $y$, $f A \neq 0$, and $H$ is a smooth function of the variables $w$ and $x$.

The formula (1.6) can be now written in the form

$$
\begin{equation*}
((k-\tilde{c}) f A)_{y}^{\prime}=0, \quad \text { i.e., } \quad k-\tilde{c}=\frac{\sigma}{f A} \tag{1.9}
\end{equation*}
$$

for some function $\sigma=\sigma(w, x) \neq 0$.
Now, define the function $\chi=\chi(w, x, y)$ of the variables $w, x$ and $y$ by

$$
\begin{equation*}
\chi=\frac{1}{f A}=\frac{k-\tilde{c}}{\sigma} \tag{1.10}
\end{equation*}
$$

Then, using (1.8) and (1.10), we obtain easily the following expression for the components of the connection form:

$$
\left\{\begin{array}{l}
\omega_{2}^{1}=-A \alpha d x+R d w+\beta d y  \tag{1.11}\\
\omega_{3}^{1}=A \beta d x+S d w \\
\omega_{3}^{2}=A_{y}^{\prime} d x+T d w
\end{array}\right.
$$

where

$$
\left\{\begin{align*}
\alpha & =\chi\left(A_{w}^{\prime}-C_{x}^{\prime}-H A_{y}^{\prime}\right)  \tag{1.12}\\
\beta & =\frac{\chi}{2}\left(H_{x}^{\prime}+A C_{y}^{\prime}-C A_{y}^{\prime}\right)
\end{align*}\right.
$$

and

$$
\left\{\begin{array}{l}
R=\chi f f_{x}^{\prime}-C \alpha+H \beta  \tag{1.13}\\
S=f_{y}^{\prime}+C \beta \\
T=C_{y}^{\prime}-f \beta
\end{array}\right.
$$

The curvature conditions (1.4) then give a system of nine partial differential equations for our problem:

$$
\begin{align*}
(A \alpha)_{y}^{\prime}+\beta_{x}^{\prime} & =0  \tag{A1}\\
R_{y}^{\prime}-\beta_{w}^{\prime} & =0
\end{align*}
$$

$$
\begin{equation*}
(A \alpha)_{w}^{\prime}+R_{x}^{\prime}+S A_{y}^{\prime}-A \beta T=-f A k \tag{A3}
\end{equation*}
$$

$$
\begin{equation*}
A_{y y}^{\prime \prime}-A \beta^{2}=-\tilde{c} A \tag{B1}
\end{equation*}
$$

$$
-A_{y w}^{\prime \prime}+T_{x}^{\prime}+A(\beta R+\alpha S)=\tilde{c} A H
$$

$$
\begin{equation*}
T_{y}^{\prime}-S \beta=-\tilde{c} C \tag{B3}
\end{equation*}
$$

$$
\begin{equation*}
(A \beta)_{y}^{\prime}+A_{y}^{\prime} \beta=0 \tag{C2}
\end{equation*}
$$

$$
\begin{equation*}
S_{y}^{\prime}+T \beta=-\tilde{c} f \tag{C3}
\end{equation*}
$$

## 2 - The first integrals and the reduction of the basic system of partial differential equations

The aim of this section is to replace the partial differential equations of the series (B) and (C) by a system of algebraic equations for the new functions depending only on the variables $w$ and $x$.

First of all, we can eliminate (B2) and (C2) by the same procedure as in [10]: the equation (B2) is a consequence of (A1) and (B1); the equation (C2) is a consequence of (A1), (A2) and (C1). Moreover, proposition 2.3 from [10] still holds (with a slight change of the notation). We have

Proposition 2.1. The equations (B3) and (C3) are satisfied if and only if

$$
\begin{equation*}
f T-C S=\varphi_{0} \tag{2.1}
\end{equation*}
$$

where $\varphi_{0}=\varphi_{0}(w, x)$ is an arbitrary function of the variables $w$ and $x$. Moreover, we have, in the hyperbolic case $\tilde{c}=-\lambda^{2}$,

$$
\begin{equation*}
S^{2}+T^{2}=\lambda\left[\varphi_{1} \cosh (2 \lambda y)+\varphi_{2} \sinh (2 \lambda y)-\varphi_{3}\right] \tag{2.2h}
\end{equation*}
$$

$$
\begin{equation*}
f S+C T=\varphi_{2} \cosh (2 \lambda y)+\varphi_{1} \sinh (2 \lambda y) \tag{2.3h}
\end{equation*}
$$

$$
\begin{equation*}
f^{2}+C^{2}=\frac{1}{\lambda}\left[\varphi_{1} \cosh (2 \lambda y)+\varphi_{2} \sinh (2 \lambda y)+\varphi_{3}\right] \tag{2.4h}
\end{equation*}
$$

where the functions $\varphi_{i}=\varphi_{i}(w, x), i=1,2,3$, of the variables $w$ and $x$ satisfy the single relation

$$
\begin{equation*}
\varphi_{0}^{2}+\varphi_{2}^{2}-\left(\varphi_{1}^{2}-\varphi_{3}^{2}\right)=0 \tag{2.5h}
\end{equation*}
$$

and in the elliptic case $\tilde{c}=\lambda^{2}$,

$$
\begin{equation*}
S^{2}+T^{2}=\lambda\left[\varphi_{1} \cos (2 \lambda y)-\varphi_{2} \sin (2 \lambda y)+\varphi_{3}\right] \tag{2.2e}
\end{equation*}
$$

$$
\begin{equation*}
f S+C T=\varphi_{2} \cos (2 \lambda y)+\varphi_{1} \sin (2 \lambda y) \tag{2.3e}
\end{equation*}
$$

$$
\begin{equation*}
f^{2}+C^{2}=\frac{1}{\lambda}\left[-\varphi_{1} \cos (2 \lambda y)+\varphi_{2} \sin (2 \lambda y)+\varphi_{3}\right] \tag{2.4e}
\end{equation*}
$$

where the functions $\varphi_{i}=\varphi_{i}(w, x), i=1,2,3$, of the variables $w$ and $x$ satisfy the single relation

$$
\begin{equation*}
\varphi_{0}^{2}+\varphi_{2}^{2}+\varphi_{1}^{2}-\varphi_{3}^{2}=0 \tag{2.5e}
\end{equation*}
$$

Proposition 2.2. From the equations (A1), (A2), (B1), (C1) and (C3), we have, in the hyperbolic case,

$$
\begin{equation*}
f A=f_{1} \cosh (2 \lambda y)+f_{2} \sinh (2 \lambda y)+f_{3} \tag{2.6h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
f A=f_{1} \cos (2 \lambda y)+f_{2} \sin (2 \lambda y)+f_{3} \tag{2.6e}
\end{equation*}
$$

where $f_{i}=f_{i}(w, x), i=1,2,3$, are some functions of the variables $w$ and $x$.

There is a function $\varphi_{4}=\varphi_{4}(w, x)$ of the variables $w$ and $x$ such that, in the hyperbolic case,

$$
\begin{equation*}
S A=\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4} \tag{2.7h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
S A=\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4} \tag{2.7e}
\end{equation*}
$$

Further, the equation (A3) is reduced to the equation

$$
\begin{equation*}
(A \alpha)_{w}^{\prime}+R_{x}^{\prime}+\tau=0 \tag{2.8}
\end{equation*}
$$

where

$$
\begin{equation*}
\tau=(S A)_{y}^{\prime}+f A \rho_{1} \tag{2.9}
\end{equation*}
$$

is a function of the variables $w$ and $x$.
Proof. From (C3) we obtain, using also (1.13),

$$
\begin{align*}
(S A)_{y}^{\prime} & =S A_{y}^{\prime}-A \beta T-\tilde{c} f A=  \tag{2.10}\\
& =f_{y}^{\prime} A_{y}^{\prime}+\beta\left(C A_{y}^{\prime}-A C_{y}^{\prime}\right)+f\left(A \beta^{2}-\tilde{c} A\right)
\end{align*}
$$

Due to (B1) we obtain

$$
\begin{align*}
(S A)_{y}^{\prime} & =f_{y}^{\prime} A_{y}^{\prime}+A_{y y}^{\prime \prime} f+\beta\left(C A_{y}^{\prime}-A C_{y}^{\prime}\right)= \\
& =\left(A_{y}^{\prime} f\right)_{y}^{\prime}+\beta\left(C A_{y}^{\prime}-A C_{y}^{\prime}\right) \tag{2.11}
\end{align*}
$$

On the other hand, using (1.13) first and (C1) later, we get

$$
\begin{equation*}
(S A)_{y}^{\prime}=\left[f_{y}^{\prime} A+(A \beta) C\right]_{y}^{\prime}=\left(f_{y}^{\prime} A\right)_{y}^{\prime}-\beta\left(C A_{y}^{\prime}-A C_{y}^{\prime}\right) . \tag{2.12}
\end{equation*}
$$

As the sum of (2.11) and (2.12) we obtain

$$
\begin{equation*}
2(S A)_{y}^{\prime}=\left(f A_{y}^{\prime}\right)_{y}^{\prime}+\left(f_{y}^{\prime} A\right)_{y}^{\prime}=(f A)_{y y}^{\prime \prime} . \tag{2.13}
\end{equation*}
$$

Using (A1) and (A2), we obtain

$$
\begin{equation*}
\left[(A \alpha)_{w}^{\prime}+R_{x}^{\prime}\right]_{y}^{\prime}=0 \tag{2.14}
\end{equation*}
$$

Due to (2.10), (1.10) and $\rho_{1}=k+\tilde{c}$, the equation (A3) takes in the form

$$
\begin{equation*}
(A \alpha)_{w}^{\prime}+R_{x}^{\prime}+(S A)_{y}^{\prime}+f A \rho_{1}=0 \tag{2.15}
\end{equation*}
$$

According to (2.14), the function $\tau$ defined by (2.9) does not depend on $y$. This together with (2.15) implies (2.8). Also, the equations (2.13) and (2.9) imply

$$
\begin{equation*}
(f A)_{y y}^{\prime \prime}+2 f A \rho_{1}=2 \tau . \tag{2.16}
\end{equation*}
$$

Substituting (1.10) and $\rho_{1}=k+\tilde{c}$ into (2.16), we obtain

$$
\begin{equation*}
\left(\frac{\sigma}{k-\tilde{c}}\right)_{y y}^{\prime \prime}+\frac{2(k+\tilde{c}) \sigma}{k-\tilde{c}}-2 \tau=0 . \tag{2.17}
\end{equation*}
$$

Because $\sigma$ does not depend on $y$, putting

$$
\begin{equation*}
F=\frac{1}{k-\tilde{c}}-\frac{\tau-\sigma}{2 \tilde{c} \sigma}, \tag{2.18}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
F_{y y}^{\prime \prime}+4 \tilde{c} F=0 . \tag{2.19}
\end{equation*}
$$

Moreover we get, from (2.18) and (1.10),

$$
\begin{equation*}
f A=F \sigma+f_{3}, \tag{2.20}
\end{equation*}
$$

where $f_{3}=f_{3}(w, x)$ is an arbitrary function of the variables $w$ and $x$.
The general solution of the partial differential equation (2.19) is, in the hyperbolic case,

$$
\begin{equation*}
F=F_{1} \cosh (2 \lambda y)+F_{2} \sinh (2 \lambda y) \tag{2.21h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
F=F_{1} \cos (2 \lambda y)+F_{2} \sin (2 \lambda y) \tag{2.21e}
\end{equation*}
$$

where $F_{1}=F_{1}(w, x)$ and $F_{2}=F_{2}(w, x)$ are arbitrary functions of the variables $w$ and $x$. This together with (2.20) implies (2.6h) and (2.6e).

From (2.6he) and (2.13) we obtain (2.7he), respectively.

Proposition 2.3. The equations (B1) and (C1) are satisfied if and only if

$$
\begin{equation*}
\beta A^{2}=\lambda a_{0} \tag{2.22}
\end{equation*}
$$

where $a_{0}=a_{0}(w, x)$ is an arbitrary function and, moreover, we have (a) in the hyperbolic case,

$$
\begin{equation*}
A^{2}=a_{1} \cosh (2 \lambda y)+a_{2} \sinh (2 \lambda y)+a_{3} \tag{2.23h}
\end{equation*}
$$

where $a_{i}=a_{i}(w, x), i=1,2,3$, are functions of the variables $w$ and $x$ satisfying

$$
\begin{equation*}
a_{0}^{2}+a_{2}^{2}-\left(a_{1}^{2}-a_{3}^{2}\right)=0 \tag{2.24h}
\end{equation*}
$$

(b) in the elliptic case,

$$
\begin{equation*}
A^{2}=a_{1} \cos (2 \lambda y)+a_{2} \sin (2 \lambda y)+a_{3} \tag{2.23e}
\end{equation*}
$$

where $a_{i}=a_{i}(w, x), i=1,2,3$, are functions of the variables $w$ and $x$ satisfying

$$
\begin{equation*}
a_{0}^{2}+a_{2}^{2}+a_{1}^{2}-a_{3}^{2}=0 \tag{2.24e}
\end{equation*}
$$

The proof is the same as for proposition 2.5 in [10] (with a slight change of the notation).

Proposition 2.4. We have, in the hyperbolic case,

$$
\begin{align*}
2 \lambda a_{0} A C= & {\left[a_{1} \varphi_{5}+2 \lambda\left(a_{2} f_{3}-a_{3} f_{2}\right)\right] \cosh (2 \lambda y)+} \\
& +\left[a_{2} \varphi_{5}-2 \lambda\left(a_{3} f_{1}-a_{1} f_{3}\right)\right] \sinh (2 \lambda y)+  \tag{2.25~h}\\
& +a_{3} \varphi_{5}-2 \lambda\left(a_{2} f_{1}-a_{1} f_{2}\right)
\end{align*}
$$

and, in the elliptic case,

$$
\begin{align*}
2 \lambda a_{0} A C= & {\left[a_{1} \varphi_{5}+2 \lambda\left(a_{2} f_{3}-a_{3} f_{2}\right)\right] \cos (2 \lambda y)+} \\
& +\left[a_{2} \varphi_{5}+2 \lambda\left(a_{3} f_{1}-a_{1} f_{3}\right)\right] \sin (2 \lambda y)+  \tag{2.25e}\\
& +a_{3} \varphi_{5}+2 \lambda\left(a_{2} f_{1}-a_{1} f_{2}\right)
\end{align*}
$$

where $\varphi_{5}=\varphi_{5}(w, x)$ is an arbitrary function of the variables $w$ and $x$.

Proof. Subtracting equations (2.11) and (2.12), we get

$$
\left(f A_{y}^{\prime}-f_{y}^{\prime} A\right)_{y}^{\prime}+2 \beta\left(A_{y}^{\prime} C-A C_{y}^{\prime}\right)=0
$$

that is,

$$
\left(f A_{y}^{\prime}-f_{y}^{\prime} A\right)_{y}^{\prime}=2 \beta A^{2} \frac{A C_{y}^{\prime}-A_{y}^{\prime} C}{A^{2}}
$$

Using (2.22), we get

$$
\begin{equation*}
\left(f A_{y}^{\prime}-f_{y}^{\prime} A\right)_{y}^{\prime}=2 \lambda a_{0}\left(\frac{C}{A}\right)_{y}^{\prime} \tag{2.26}
\end{equation*}
$$

Integrating (2.26) with respect to $y$ and multiplying by $A^{3}$, we get

$$
\begin{equation*}
2 \lambda a_{0} A C=\varphi_{5} A^{2}+(f A)\left(A^{2}\right)_{y}^{\prime}-A^{2}(f A)_{y}^{\prime} \tag{2.27}
\end{equation*}
$$

where $\varphi_{5}=\varphi_{5}(w, x)$ is an arbitrary function of the variables $w$ and $x$. Substituting (2.6he) and (2.23he) into (2.27), we obtain our assertions, respectively.

The following proposition is more explicit.
Proposition 2.5. We have, in the hyperbolic case,

$$
\begin{equation*}
A C=b_{1} \cosh (2 \lambda y)+b_{2} \sinh (2 \lambda y)+b_{3} \tag{2.28h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
A C=b_{1} \cos (2 \lambda y)+b_{2} \sin (2 \lambda y)+b_{3} \tag{2.28e}
\end{equation*}
$$

where $b_{i}=b_{i}(w, x), i=1,2,3$, are functions of the variables $w$ and $x$.
Proof. For $a_{0} \neq 0$, the assertion (2.28he) is a direct consequence of (2.25he), respectively.

Suppose now $\tilde{c}=\epsilon \lambda^{2}, \epsilon= \pm 1$, and $a_{0}=0$. Then $\beta=0$ by (2.22) and we get from $(1.13)_{3}$ and (B3) that

$$
C_{y y}^{\prime \prime}=-\tilde{c} C=-\epsilon \lambda^{2} C
$$

Hence we get, in the hyperbolic case,

$$
\begin{equation*}
C=r \cosh (\lambda y)+s \sinh (\lambda y) \tag{2.29h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
C=r \cos (\lambda y)+s \sin (\lambda y) \tag{2.29e}
\end{equation*}
$$

where $r=r(w, x)$ and $s=s(w, x)$ are arbitrary functions of the variables $w$ and $x$. On the other hand, (2.23he) and (2.24he) with $a_{0}=0$ imply, in the hyperbolic case,

$$
\begin{equation*}
A=p \cosh (\lambda y)+q \sinh (\lambda y) \tag{2.30h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
A=p \cos (\lambda y)+q \sin (\lambda y) \tag{2.30e}
\end{equation*}
$$

with some functions $p=p(w, x)$ and $q=q(w, x)$ of the variables $w$ and $x$. Hence (2.28he) follows.

Remark. We denote $\operatorname{sgn} \tilde{c}$ by $\epsilon$ in the sequel. This notation will be used later to unify many formulas for the hyperbolic and the elliptic case.

Now we introduce the function $h=h(w, x)$ by

$$
\begin{equation*}
h=H_{x}^{\prime} \tag{2.31}
\end{equation*}
$$

Proposition 2.6. We have

$$
\left\{\begin{align*}
h a_{1} & =2 \lambda\left[a_{0} f_{1}+a_{2} b_{3}-a_{3} b_{2}\right],  \tag{2.32}\\
h a_{2} & =2 \lambda\left[a_{0} f_{2}+\epsilon\left(a_{3} b_{1}-a_{1} b_{3}\right)\right] \\
h a_{3} & =2 \lambda\left[a_{0} f_{3}-\left(a_{1} b_{2}-a_{2} b_{1}\right)\right]
\end{align*}\right.
$$

Proof. From (1.12) ${ }_{2}$ we get

$$
h=2 f A \beta-(A C)_{y}^{\prime}+2 A_{y}^{\prime} C
$$

Then (2.22) and (1.10) imply

$$
\begin{equation*}
h A^{2}=2 \lambda a_{0} f A-A^{2}(A C)_{y}^{\prime}+(A C)\left(A^{2}\right)_{y}^{\prime} \tag{2.33}
\end{equation*}
$$

Now we use (2.6he), (2.23he) and (2.28he) to get (2.32he).
From (2.21he), (1.10) and (2.1) we obtain

$$
\begin{equation*}
S=f \chi Q, \quad T=C \chi Q+\varphi_{0} \chi A \tag{2.34}
\end{equation*}
$$

where, in the hyperbolic case,

$$
\begin{equation*}
Q=\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4} \tag{2.35h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
Q=\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4} \tag{2.35e}
\end{equation*}
$$

Substituting from (2.34) into the partial differential equation (C3), we obtain, using also (2.22),

$$
\left(f \chi Q_{y}^{\prime}-\frac{A_{y}^{\prime}}{A^{2}} Q\right) A^{2}+\lambda a_{0} C \chi Q+\lambda a_{0} \varphi_{0} \chi A=-\tilde{c} f A^{2}
$$

Multiplying this equation by $A$ and using (2.27) and (1.10), we get

$$
\begin{equation*}
2 f A Q_{y}^{\prime}+\varphi_{5} Q-Q(f A)_{y}^{\prime}+2 \lambda a_{0} \varphi_{0}+2 \tilde{c}(f A)^{2}=0 \tag{2.36}
\end{equation*}
$$

Substituting from (2.6he) and (2.35he) into (2.36), we obtain

$$
\left\{\begin{array}{l}
f_{1}\left(\varphi_{5}-2 \varphi_{4}\right)=0, \quad f_{2}\left(\varphi_{5}-2 \varphi_{4}\right)=0  \tag{2.37}\\
\varphi_{4} \varphi_{5}+2 \lambda a_{0} \varphi_{0}-2 \lambda^{2}\left[f_{2}^{2}+\epsilon\left(f_{1}^{2}-f_{3}^{2}\right)\right]=0
\end{array}\right.
$$

Substituting (2.35he) into (2.34), we obtain, in the hyperbolic case,
(2.38h) $S=f \chi\left[\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4}\right]$,

$$
\begin{equation*}
T=C \chi\left[\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4}\right]+\varphi_{0} \chi A \tag{2.39h}
\end{equation*}
$$

and, in the elliptic case,

$$
\begin{equation*}
S=f \chi\left[\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4}\right] \tag{2.38e}
\end{equation*}
$$

(2.39e) $\quad T=C \chi\left[\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4}\right]+\varphi_{0} \chi A$.

Hence we obtain, in the hyperbolic case,

$$
\begin{align*}
& f A(C T+f S)= \\
& =\varphi_{0} A C+\left[\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4}\right]\left(f^{2}+C^{2}\right) \tag{2.40h}
\end{align*}
$$

and, in the elliptic case,

$$
\begin{align*}
& f A(C T+f S)= \\
& =\varphi_{0} A C+\left[\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4}\right]\left(f^{2}+C^{2}\right) \tag{2.40e}
\end{align*}
$$

Substituting (2.3he), (2.4he) and (2.6he) into (2.40he), we get in the hyperbolic case,

$$
\begin{align*}
\varphi_{0} A C= & \left(f_{3} \varphi_{2}-f_{2} \varphi_{3}-\frac{1}{\lambda} \varphi_{1} \varphi_{4}\right) \cosh (2 \lambda y)+ \\
& +\left(f_{3} \varphi_{1}-f_{1} \varphi_{3}-\frac{1}{\lambda} \varphi_{2} \varphi_{4}\right) \sinh (2 \lambda y)+  \tag{2.41h}\\
& +f_{1} \varphi_{2}-f_{2} \varphi_{1}-\frac{1}{\lambda} \varphi_{3} \varphi_{4}
\end{align*}
$$

and, in the elliptic case,

$$
\begin{align*}
\varphi_{0} A C= & \left(f_{3} \varphi_{2}-f_{2} \varphi_{3}+\frac{1}{\lambda} \varphi_{1} \varphi_{4}\right) \cos (2 \lambda y)+ \\
& +\left(f_{3} \varphi_{1}+f_{1} \varphi_{3}-\frac{1}{\lambda} \varphi_{2} \varphi_{4}\right) \sin (2 \lambda y)+  \tag{2.41e}\\
& +f_{1} \varphi_{2}+f_{2} \varphi_{1}-\frac{1}{\lambda} \varphi_{3} \varphi_{4}
\end{align*}
$$

Another consequence of (2.38he) and (2.39he) is, in the hyperbolic case,

$$
\begin{align*}
& (f A)^{2}\left(S^{2}+T^{2}\right)=\left[\lambda^{2} f_{2}^{2} \cosh ^{2}(2 \lambda y)+\right. \\
& \quad+\lambda^{2} f_{1}^{2} \sinh ^{2}(2 \lambda y)+2 \lambda^{2} f_{1} f_{2} \cosh (2 \lambda y) \sinh (2 \lambda y)+ \\
& \left.\quad+2 \lambda f_{2} \varphi_{4} \cosh (2 \lambda y)+2 \lambda f_{1} \varphi_{4} \sinh (2 \lambda y)+\varphi_{4}^{2}\right]\left(f^{2}+C^{2}\right)+  \tag{2.42~h}\\
& \quad+2 \varphi_{0} A C\left[\lambda f_{2} \cosh (2 \lambda y)+\lambda f_{1} \sinh (2 \lambda y)+\varphi_{4}\right]+\varphi_{0}^{2} A^{2}
\end{align*}
$$

and, in the elliptic case,

$$
\begin{align*}
& (f A)^{2}\left(S^{2}+T^{2}\right)=\left[\lambda^{2} f_{2}{ }^{2} \cos ^{2}(2 \lambda y)+\right. \\
& \quad+\lambda^{2} f_{1}{ }^{2} \sin ^{2}(2 \lambda y)-2 \lambda^{2} f_{1} f_{2} \cos (2 \lambda y) \sin (2 \lambda y)+ \\
& \left.\quad+2 \lambda f_{2} \varphi_{4} \cos (2 \lambda y)-2 \lambda f_{1} \varphi_{4} \sin (2 \lambda y)+\varphi_{4}{ }^{2}\right]\left(f^{2}+C^{2}\right)+  \tag{2.42e}\\
& \quad+2 \varphi_{0} A C\left[\lambda f_{2} \cos (2 \lambda y)-\lambda f_{1} \sin (2 \lambda y)+\varphi_{4}\right]+\varphi_{0}^{2} A^{2}
\end{align*}
$$

Using the formulas (2.2he), (2.4he), (2.6he), (2.23he) and (2.41he), we obtain from (2.42he)

$$
\left\{\begin{align*}
& \lambda \varphi_{0}^{2} a_{1}=\varphi_{1}[ \left.\lambda^{2}\left(f_{1}^{2}-\epsilon f_{2}^{2}+f_{3}^{2}\right)-\epsilon \varphi_{4}^{2}\right]+  \tag{2.43}\\
&+2 \lambda^{2} f_{1}\left(\epsilon f_{3} \varphi_{3}-f_{2} \varphi_{2}\right)+2 \lambda \varphi_{4}\left(f_{2} \varphi_{3}-f_{3} \varphi_{2}\right) \\
& \lambda \varphi_{0}^{2} a_{2}=\epsilon \varphi_{2}\left[\lambda^{2}\left(f_{1}^{2}-\epsilon f_{2}^{2}-f_{3}^{2}\right)+\epsilon \varphi_{4}^{2}\right]+ \\
&+2 \lambda^{2} f_{2}\left(f_{1} \varphi_{1}+\epsilon f_{3} \varphi_{3}\right)-2 \lambda \varphi_{4}\left(f_{3} \varphi_{1}+\epsilon f_{1} \varphi_{3}\right) \\
& \lambda \varphi_{0}^{2} a_{3}=\epsilon \varphi_{3}\left[\lambda^{2}\left(f_{1}^{2}+\epsilon f_{2}^{2}+f_{3}^{2}\right)+\epsilon \varphi_{4}^{2}\right]+ \\
&+2 \lambda^{2} f_{3}\left(f_{1} \varphi_{1}-f_{2} \varphi_{2}\right)-2 \lambda \varphi_{4}\left(f_{1} \varphi_{2}+\epsilon f_{2} \varphi_{1}\right)
\end{align*}\right.
$$

Consider now the identity $(A C)^{2}=A^{2}\left(f^{2}+C^{2}\right)-(A f)^{2}$. Substituting from (2.4he), (2.6he), (2.23he) and (2.28he), we get a system of quadratic
equations

$$
\left\{\begin{array}{l}
\lambda\left(b_{1}^{2}-\epsilon b_{2}^{2}+f_{1}^{2}-\epsilon f_{2}^{2}\right)=-\epsilon\left(a_{1} \varphi_{1}+a_{2} \varphi_{2}\right)  \tag{2.44}\\
\lambda\left(b_{1}^{2}+\epsilon b_{2}^{2}+2 b_{3}^{2}+f_{1}^{2}+\epsilon f_{2}^{2}+2 f_{3}^{2}\right)=-\epsilon\left(a_{1} \varphi_{1}-a_{2} \varphi_{2}\right)+2 a_{3} \varphi_{3} \\
2 \lambda\left(b_{1} b_{2}+f_{1} f_{2}\right)=a_{1} \varphi_{2}-\epsilon a_{2} \varphi_{1} \\
2 \lambda\left(b_{1} b_{3}+f_{1} f_{3}\right)=a_{1} \varphi_{3}-\epsilon a_{3} \varphi_{1} \\
2 \lambda\left(b_{2} b_{3}+f_{2} f_{3}\right)=a_{2} \varphi_{3}+a_{3} \varphi_{2}
\end{array}\right.
$$

In the notation (2.28he) we can rewrite (2.25he) in the form

$$
\left\{\begin{array}{l}
2 \lambda a_{0} b_{1}=a_{1} \varphi_{5}+2 \lambda\left(a_{2} f_{3}-a_{3} f_{2}\right)  \tag{2.45}\\
2 \lambda a_{0} b_{2}=a_{2} \varphi_{5}+2 \epsilon \lambda\left(a_{3} f_{1}-a_{1} f_{3}\right) \\
2 \lambda a_{0} b_{3}=a_{3} \varphi_{5}-2 \lambda\left(a_{1} f_{2}-a_{2} f_{1}\right)
\end{array}\right.
$$

Also, we can rewrite (2.41he) in the form

$$
\left\{\begin{array}{l}
\lambda \varphi_{0} b_{1}=-\lambda\left(f_{2} \varphi_{3}-f_{3} \varphi_{2}\right)+\epsilon \varphi_{1} \varphi_{4},  \tag{2.46}\\
\lambda \varphi_{0} b_{2}=\lambda\left(f_{3} \varphi_{1}+\epsilon f_{1} \varphi_{3}\right)-\varphi_{2} \varphi_{4}, \\
\lambda \varphi_{0} b_{3}=\lambda\left(f_{1} \varphi_{2}+\epsilon f_{2} \varphi_{1}\right)-\varphi_{3} \varphi_{4} .
\end{array}\right.
$$

Proposition 2.7. If $a_{0} \neq 0$, then we have

$$
\begin{equation*}
h=-\frac{2 \lambda\left[\epsilon\left(a_{1} f_{1}-a_{3} f_{3}\right)+a_{2} f_{2}\right]}{a_{0}} \tag{2.47}
\end{equation*}
$$

Proof. The assertion follows from (2.32), (2.45) and (2.24he).
Now we have the main results of this section.
THEOREM 2.8. Let $\lambda$ be a nonzero constant. Let $\varphi_{0}, \varphi_{1}, \ldots, \varphi_{5}, a_{0}$, $a_{1}, a_{2}, a_{3}, b_{1}, b_{2}, b_{3}, f_{1}, f_{2}, f_{3}$ and $h$ be functions of two variables $w$ and $x$ defined in some domain $V \subset \mathbb{R}^{2}(w, x)$, satisfying eight collections of algebraic equations (2.5), (2.24), (2.32), (2.37) $)_{2},(2.43),(2.44),(2.45)$ and (2.46) (either of hyperbolic type, or of elliptic type) with the corresponding parameter $\lambda$, and such that $a_{1}{ }^{2}+{a_{2}}^{2}+a_{3}{ }^{2}>0$ in $V$.

Let $A, f, C$ and $H$ be functions defined in a domain $U \subset \mathbb{R}^{3}(w, x, y)$, where $A \neq 0$, by the formulas (2.23), (2.6), (2.28) and (2.31) of the corresponding type, and let the metric $g$ be defined on $U$ by (1.8). Further, let $\alpha, \beta$ and $R$ be defined as in $(1.12)_{1},(2.22),(1.13)_{1}$. Then the curvature conditions (1.4) are satisfied for some function $k=k(w, x, y)$ of the variables $w, x$ and $y$, and for the corresponding constant $\tilde{c}= \pm \lambda^{2}$ if and only if the system of partial differential equations (A1) and (A2) is satisfied.

Proof. The assertion follows from the whole series of propositions and formulas given in this section. The only point here is to show that, if we do not prescribe the function $k=k(w, x, y)$ in advance, then the equation (A3) (or, equivalently, (2.8)) does not give any additional condition. But, due to (2.9) and (1.2), the equation (2.8) can be considered just as a formula for calculating the Ricci eigenvalue $\rho_{1}$ or the scalar curvature $\mathrm{Sc}(g)=2 k+4 \tilde{c}$ of $(M, g)$.

Remark. The algebraic conditions mentioned above are, of course, far from being independent, but they are all useful.

We conclude this section by proving additional algebraic equations between our basic functions.

Proposition 2.9. We have

$$
\begin{equation*}
\varphi_{5}=2 \varphi_{4} \tag{2.48}
\end{equation*}
$$

$$
\begin{equation*}
\varphi_{0} A^{2}-\lambda a_{0}\left(f^{2}+C^{2}\right)+\varphi_{5} A C+h f A=0 \tag{2.49}
\end{equation*}
$$

Proof. If $f_{1}^{2}+f_{2}^{2} \neq 0$, then (2.48) follows from (2.37) $)_{1,2}$. If $f_{1}=$ $f_{2}=0$, then we proceed as in the proof of proposition 4.1 in [10].

To derive (2.49), we rewrite (2.37) using (2.48) in the form

$$
\begin{equation*}
\lambda a_{0} \varphi_{0}=\lambda^{2}\left[f_{2}^{2}+\epsilon\left({f_{1}}^{2}-{f_{3}}^{2}\right)\right]-\varphi_{4}^{2} . \tag{2.50}
\end{equation*}
$$

Suppose $a_{0} \neq 0$. Then (2.45) and (2.48) imply

$$
\left\{\begin{align*}
b_{1} & =\frac{a_{1} \varphi_{4}+\lambda\left(a_{2} f_{3}-a_{3} f_{2}\right)}{\lambda a_{0}}  \tag{2.51}\\
b_{2} & =\frac{a_{2} \varphi_{4}+\epsilon \lambda\left(a_{3} f_{1}-a_{1} f_{3}\right)}{\lambda a_{0}} \\
b_{3} & =\frac{a_{3} \varphi_{4}-\lambda\left(a_{1} f_{2}-a_{2} f_{1}\right)}{\lambda a_{0}}
\end{align*}\right.
$$

Now we substitute for $A^{2}, f^{2}+C^{2}, A C, \varphi_{5}, h$ and $f A$ of the left hand side of (2.49) from (2.23he), (2.4he), (2.28he), (2.48) and (2.6he), respectively. Then the identity (2.49) follows. If $a_{0}=0$, we use the direct check as in [10].

Proposition 2.10. The following algebraic formulas hold

$$
\begin{align*}
& 2 \lambda\left(a_{1} f_{1}+\epsilon a_{2} f_{2}-a_{3} f_{3}\right)=-\epsilon a_{0} h  \tag{2.52}\\
& 4 \lambda^{2}\left(b_{1} f_{1}+\epsilon b_{2} f_{2}-b_{3} f_{3}\right)=-\epsilon \varphi_{5} h \\
& 2 \lambda\left(\varphi_{1} f_{1}-\varphi_{2} f_{2}-\varphi_{3} f_{3}\right)=\epsilon \varphi_{0} h \\
& 2 \lambda\left(a_{1} b_{1}+\epsilon a_{2} b_{2}-a_{3} b_{3}\right)=-\epsilon a_{0} \varphi_{5}
\end{align*}
$$

Proof. From (2.24he) and (2.32) we obtain

$$
2 \lambda a_{0}\left(a_{1} f_{1}+\epsilon a_{2} f_{2}-a_{3} f_{3}\right)=-\epsilon a_{0}^{2} h
$$

Hence we obtain (2.52) if $a_{0} \neq 0$. From (2.45) and (2.24he) we obtain

$$
2 \lambda a_{0}\left(b_{1} f_{1}+\epsilon b_{2} f_{2}-b_{3} f_{3}\right)=\varphi_{5}\left(a_{1} f_{1}+\epsilon a_{2} f_{2}-a_{3} f_{3}\right)
$$

which together with (2.52) implies (2.53) when $a_{0} \neq 0$. From (2.46) we obtain

$$
\varphi_{4}\left(\varphi_{1} f_{1}-\varphi_{2} f_{2}-\varphi_{3} f_{3}\right)=-\lambda \varphi_{0}\left(b_{1} f_{1}+b_{2} f_{2}-b_{3} f_{3}\right)
$$

hence, if $a_{0} \varphi_{4} \neq 0$, we obtain (2.54) using (2.53) and (2.48). Finally from (2.45) we obtain

$$
2 \lambda a_{0}\left(a_{1} b_{1}+\epsilon a_{2} b_{2}-a_{3} b_{3}\right)=-\epsilon a_{0}^{2} \varphi_{5}
$$

Thus we obtain (2.55) when $a_{0} \varphi_{4} \neq 0$.
For $a_{0} \varphi_{4}=0$ we use the continuity argument or a rather lengthy direct check (cf. proposition 4.10 in [9]).

## 3 - The Riemannian invariants

Let $(M, g)$ be given locally as in proposition 1.1. We rewrite the formulas (1.11) using the forms $\omega^{1}, \omega^{2}$ and $\omega^{3}$ as a basis. It follows

$$
\left\{\begin{array}{l}
\omega_{2}^{1}=\chi f_{x}^{\prime} \omega^{1}-\alpha \omega^{2}+\beta \omega^{3},  \tag{3.1}\\
\omega_{3}^{1}=\frac{f_{y}^{\prime}}{f} \omega^{1}+\beta \omega^{2} \\
\omega_{3}^{2}=(\beta-h \chi) \omega^{1}+\frac{A_{y}^{\prime}}{A} \omega^{2}, \quad h=H_{x}^{\prime}
\end{array}\right.
$$

We also write, for brevity,

$$
\begin{equation*}
\omega_{3}^{1}=a \omega^{1}+b \omega^{2}, \quad \omega_{3}^{2}=c \omega^{1}+e \omega^{2} \tag{3.2}
\end{equation*}
$$

where

$$
\begin{equation*}
a=\frac{f_{y}^{\prime}}{f}, \quad b=\beta, \quad c=\beta-h \chi, \quad e=\frac{A_{y}^{\prime}}{A} \tag{3.3}
\end{equation*}
$$

Using the standard formula $\nabla_{E_{j}} E_{i}=\sum_{k} \omega_{i}^{k}\left(E_{j}\right) E_{k}, i, j=1,2,3$, from [7], we obtain

$$
\begin{cases}\nabla_{E_{1}} E_{1}=-\chi f_{x}^{\prime} E_{2}-a E_{3}, & \nabla_{E_{1}} E_{2}=\chi f_{x}^{\prime} E_{1}-c E_{3}  \tag{3.4}\\ \nabla_{E_{2}} E_{1}=\alpha E_{2}-b E_{3}, & \nabla_{E_{2}} E_{2}=-\alpha E_{1}-e E_{3} \\ \nabla_{E_{1}} E_{3}=a E_{1}+c E_{2}, & \nabla_{E_{2}} E_{3}=b E_{1}+e E_{2} \\ \nabla_{E_{3}} E_{1}=-b E_{2}, & \nabla_{E_{3}} E_{2}=b E_{1}, \quad \nabla_{E_{3}} E_{3}=0\end{cases}
$$

The last formula shows that the trajectories of the unit vector field $E_{3}$ (consisting of the eigenvectors of the Ricci tensor $\hat{R}$ corresponding to $\left.\rho_{3}=2 \tilde{c}\right)$ are geodesics.

For the Ricci tensor $\hat{R}$ we get, using the notation (1.2) and the adapted local orthonormal coframe $\left\{\omega^{1}, \omega^{2}, \omega^{3}\right\}$,

$$
\begin{equation*}
\hat{R}=(k+\tilde{c})\left(\omega^{1} \otimes \omega^{1}+\omega^{2} \otimes \omega^{2}\right)+2 \tilde{c}\left(\omega^{3} \otimes \omega^{3}\right) . \tag{3.5}
\end{equation*}
$$

Using (3.1), (3.2) and the standard formula $\nabla_{X} \omega^{i}=-\sum_{j} \omega_{j}^{i}(X) \omega^{j}$, we obtain

$$
\begin{align*}
\nabla \hat{R}=d k & \otimes\left(\omega^{1} \otimes \omega^{1}+\omega^{2} \otimes \omega^{2}\right)+ \\
& +(\tilde{c}-k)\left\{\left(a \omega^{1}+b \omega^{2}\right) \otimes\left(\omega^{1} \otimes \omega^{3}+\omega^{3} \otimes \omega^{1}\right)+\right.  \tag{3.6}\\
& \left.+\left(c \omega^{1}+e \omega^{2}\right) \otimes\left(\omega^{2} \otimes \omega^{3}+\omega^{3} \otimes \omega^{2}\right)\right\}
\end{align*}
$$

where $a, b, c$ and $e$ are given by (3.3). Hence we also get

$$
\begin{align*}
\|\nabla \hat{R}\|^{2} & =2\|d k\|^{2}+2(\tilde{c}-k)^{2}\left(a^{2}+b^{2}+c^{2}+e^{2}\right)=  \tag{3.7}\\
& =2\left\|d \rho_{1}\right\|^{2}+2\left(\rho_{1}-\rho_{3}\right)^{2}\left(a^{2}+b^{2}+c^{2}+e^{2}\right)
\end{align*}
$$

Because $\hat{R}$ is a Riemannian invariant tensor, $\nabla \hat{R}$ is an invariant tensor. Also, because $E_{3}=\partial / \partial y$ is uniquely determined by the geometry of ( $M, g$ ) up to sign, $\omega^{3} \otimes \omega^{3}$ is an invariant tensor. Hence we see from (3.5) and (3.6) that the tensor

$$
\begin{align*}
& Q=\left(a \omega^{1}+b \omega^{2}\right) \otimes\left(\omega^{1} \otimes \omega^{3}+\omega^{3} \otimes \omega^{1}\right)+  \tag{3.8}\\
&+\left(c \omega^{1}+e \omega^{2}\right) \otimes\left(\omega^{2} \otimes \omega^{3}+\omega^{3} \otimes^{2}\right)
\end{align*}
$$

is also invariant. Now because $E_{1}$ and $E_{2}$ are determined up to an orthogonal transformation (with functional coefficients), the functions

$$
\left\{\begin{array}{l}
Q\left(E_{1}, E_{1}, E_{3}\right)+Q\left(E_{2}, E_{2}, E_{3}\right)=a+e,  \tag{3.9}\\
Q\left(E_{2}, E_{1}, E_{3}\right)-Q\left(E_{1}, E_{2}, E_{3}\right)=b-c
\end{array}\right.
$$

are Riemannian invariants up to sign.
The square of the norm $\|Q\|^{2}=2\left(a^{2}+b^{2}+c^{2}+e^{2}\right)$ is a Riemannian invariant and hence (equivalently) $a e-b c$ is a Riemannian invariant. We summarize:

Proposition 3.1. The function $a e-b c$ is a Riemannian invariant, and $a+e$ and $b-c$ are Riemannian invariants up to sign (i.e., depending on the orientation of the principal geodesics). Further, the partial derivative of any Riemannian invariant with respect to $y$ is a Riemannian invariant up to sign.

Using (1.10), we get

$$
\left\{\begin{array}{l}
a+e=(\ln (f A))_{y}^{\prime}=-\left(\ln \left(k-\epsilon \lambda^{2}\right)\right)_{y}^{\prime}  \tag{3.10}\\
b-c=h \chi=\frac{h\left(k-\epsilon \lambda^{2}\right)}{\sigma}
\end{array}\right.
$$

Further we have

$$
\begin{equation*}
a e-b c=\epsilon\left(2 \lambda^{2} f_{3} \chi-\lambda^{2}\right) \tag{3.11}
\end{equation*}
$$

The last formula is obtained by lengthy calculations using (2.52) and the obvious identities

$$
\begin{align*}
\left(A A_{y}^{\prime}\right)^{2}+\lambda^{2} a_{0}^{2} & =-\epsilon \lambda^{2}\left[\left(A^{2}-a_{3}\right)^{2}-a_{3}^{2}\right]  \tag{3.12}\\
A^{3} f_{y}^{\prime} & =(f A)_{y}^{\prime} A^{2}-(f A)\left(A A_{y}^{\prime}\right) \tag{3.13}
\end{align*}
$$

Using (3.11) we see that, in the hyperbolic case,

$$
\begin{equation*}
\frac{f A}{f_{3}}=\frac{f_{1} \cosh (2 \lambda y)+f_{2} \sinh (2 \lambda y)+f_{3}}{f_{3}} \tag{3.14h}
\end{equation*}
$$

is a Riemannian invariant and, in the elliptic case,

$$
\begin{equation*}
\frac{f A}{f_{3}}=\frac{f_{1} \cos (2 \lambda y)+f_{2} \sin (2 \lambda y)+f_{3}}{f_{3}} \tag{3.14e}
\end{equation*}
$$

is a Riemannian invariant (assuming $f_{3} \neq 0$ everywhere). (According to $(3.10)_{2}, f A / h$ and $f_{3} / h$ are then Riemannian invariants up to sign assuming $h \neq 0$ everywhere.)

Next, we give some simple results concerning isometry of Riemannian manifolds with the Ricci eigenvalues $\rho_{1}=\rho_{2}$ and nonzero constant $\rho_{3}$ to
be used later. Let $(M, g)$ be such a manifold with the metric $g$ given by (1.8) and let $(\bar{M}, \bar{g})$ be another such manifold with the metric $\bar{g}$ given by the orthonormal coframe

$$
\begin{equation*}
\bar{\omega}^{1}=\bar{f} d \bar{w}, \quad \bar{\omega}^{2}=\bar{A} d \bar{x}+\bar{C} d \bar{w}, \quad \bar{\omega}^{3}=d \bar{y}+\bar{H} d \bar{w} . \tag{3.15}
\end{equation*}
$$

Suppose that there is an isometry $\Phi:(M, g) \longrightarrow(\bar{M}, \bar{g})$ given by

$$
\begin{equation*}
\bar{w}=\bar{w}(w, x, y), \quad \bar{x}=\bar{x}(w, x, y), \quad \bar{y}=\bar{y}(w, x, y) \tag{3.16}
\end{equation*}
$$

Here we identify $\bar{w}, \bar{x}$ and $\bar{y}$ with $\bar{w} \circ \Phi, \bar{x} \circ \Phi$ and $\bar{y} \circ \Phi$, respectively.
Propositions 5.2 and 5.3 from [9] still hold without change. We have:
Proposition 3.2. The equation (3.16) can be reduced to the form

$$
\begin{equation*}
\bar{w}=\bar{w}(w, x), \quad \bar{x}=\bar{x}(w, x), \quad \bar{y}=\varepsilon y+\phi(w, x), \quad \varepsilon= \pm 1 \tag{3.17}
\end{equation*}
$$

where $\phi=\phi(w, x)$ is an arbitrary function of the variables $w$ and $x$.

## 4 - The asymptotic foliations and four types of spaces

Recall that the principal geodesics are trajectories of the vector field $E_{3}$. We introduce two basic definitions.

DEfinition 4.1. A smooth surface $N \subset(M, g)$ is called an asymptotic leaf if it is generated by the principal geodesics and its tangent planes are parallel along these principal geodesics with respect to the Riemannian connection $\nabla$ of $(M, g)$.

Definition 4.2. An asymptotic distribution on $M$ is a two-dimensional distribution which is integrable and whose integral manifolds are asymptotic leaves. The integral manifolds of an asymptotic distribution determine a foliation of $M$, which is called an asymptotic foliation.

Let $N$ be an asymptotic leaf. Then we see as in [10] that the tangent distribution of $N$ satisfies the quadratic equation

$$
\begin{equation*}
c\left(\omega^{1}\right)^{2}+(e-a) \omega^{1} \omega^{2}-b\left(\omega^{2}\right)^{2}=0 \tag{4.1}
\end{equation*}
$$

where $a, b, c$ and $e$ are given by (3.3). Of course, an asymptotic distribution must satisfy these equations locally on the whole of $M$. Conversely, any smooth distribution satisfying (4.1) is an asymptotic distribution.

The following proposition is almost obvious.
Proposition 4.3. Let $\Delta=(e-a)^{2}+4 b c$ be the discriminant of the quadratic equation (4.1). Then we have:
(E) If $\Delta<0$ on $(M, g)$, then there is no real asymptotic distribution on $M$.
(H) If $\Delta>0$ on $(M, g)$, then there are exactly two different asymptotic distributions on $M$.
(P) If $\Delta=0$ on $(M, g)$ and some of the functions $a_{0}, \varphi_{0}$ and $\varphi_{5}$ are nonzero at each point, then there is a unique asymptotic distribution on $M$.
( $\mathrm{P} \ell$ ) If $a_{0}=\varphi_{0}=\varphi_{5}=0$ on $M$, then any $\pi$-projectable smooth twodimensional distribution on $M$ is asymptotic.
(For the last part we only have to prove $b=c=e-a=0$, which follows from the formulas $(2.22),(2.27)$ and (2.49).)

Definition 4.4. A space $(M, g)$ is said to be of subtype $(\mathrm{E}),(\mathrm{H}),(\mathrm{P})$ or $(\mathrm{P} \ell)$, respectively, if the corresponding case of proposition 4.3 holds on the whole of $M$.

Let us remark that the above symbols are abbreviations for "elliptic", "hyperbolic", "parabolic" and "planar" (cf. [9]). Yet, the reader should keep in mind that, e.g., "elliptic subtype" from proposition 4.3 and "elliptic case" from proposition 2.1 are completely different geometric notions, which can be combined.

We add some more details:
Proposition 4.5. The equation (4.1) is equivalent to the equation

$$
\begin{equation*}
\lambda a_{0} d x^{2}+\varphi_{5} d x d w-\varphi_{0} d w^{2}=0 \tag{4.2}
\end{equation*}
$$

Proof. We can apply the same procedure as in [9] (proof of theorem 6.5). Here we use formulas (2.49) and (2.27) for this purpose.

Hence we can decide about the subtype of the space $(M, g)$ according to the following

Proposition 4.6. Let $\Delta^{\prime}=\varphi_{5}{ }^{2}+4 \lambda a_{0} \varphi_{0}$ be the discriminant of the quadratic equation (4.2). Then the analogue of proposition 4.3 holds if $\Delta$ is replaced by $\Delta^{\prime}$.

Proof. One can show easily that $\Delta^{\prime}=(f A)^{2} \Delta$.
Also, notice that $\Delta^{\prime}$ is given alternatively by the formula

$$
\begin{equation*}
\Delta^{\prime}=4 \lambda^{2}\left[f_{2}^{2}+\epsilon\left(f_{1}^{2}-f_{3}^{2}\right)\right] \tag{4.3}
\end{equation*}
$$

Indeed, combining $(2.37)_{3}$ with (2.48), we obtain at once

$$
\begin{equation*}
\varphi_{5}^{2}+4 \lambda a_{0} \varphi_{0}-4 \lambda^{2}\left[f_{2}^{2}+\epsilon\left(f_{1}^{2}-f_{3}^{2}\right)\right]=0 \tag{4.4}
\end{equation*}
$$

Here we stop our consideration about asymptotically foliated spaces, i.e., those of the non-elliptic subtype. These spaces have been treated in $[3$, Chapter 11 (see [14]). The present authors showed that, for each asymptotically foliated space, there is an adapted system of local coordinates such that $a_{0}=0$ identically. Then the expression for $A, C$ and $f$ are "linearized" (cf. (2.29he) and (2.30he) above) and the solution of the problem is dramatically simplified. In [14] we have shown that (generically), for each subtype $(\mathrm{H}),(\mathrm{P})$ and $(\mathrm{P} \ell)$, the general solution can be expressed in a closed form, i.e., in the form involving only arbitrary functions, algebraic operations, elementary functions, differentiations and integrations.

In the rest of this paper we concentrate ourselves only on the more complicated elliptic subtype (E).

## 5 - The quasiexplicit classification of spaces of elliptic subtype

The spaces of elliptic subtype are much more difficult to deal with because the coefficients $A, f$ and $C$ in (1.8) cannot be expressed in general in the form of linear combinations of $\cosh (\lambda y)$ and $\sinh (\lambda y)$; or of $\cos (\lambda y)$ and $\sin (\lambda y)$. We are not able to solve the classification problem explicitly, but we can still prove the local isometry classes of metrics depend
on essentially three arbitrary functions of two variables. Also, we give an example of an explicit family of metrics depending on two arbitrary functions of two variables.

We see first that the functions $a_{0}$ and $\varphi_{0}$ are always non-zero on a space of subtype (E) (cf. (4.2)). Also, we must have $h \neq 0$. (If $h=0$, then $b=c$ in (3.3) and hence $\Delta \geq 0$ in proposition 4.3.) From (2.24he) and (2.5he) we see that

$$
\begin{equation*}
\epsilon\left(a_{1}^{2}-a_{3}^{2}\right)+a_{2}^{2}<0, \quad \epsilon\left(\varphi_{1}^{2}-\varphi_{3}^{2}\right)+\varphi_{2}^{2}<0 \tag{5.1}
\end{equation*}
$$

and, from (4.3), we have

$$
\begin{equation*}
\epsilon\left(f_{1}^{2}-f_{3}^{2}\right)+f_{2}^{2}<0 \tag{5.2}
\end{equation*}
$$

We start with the following simplification:
Proposition 5.1. Every metric $g$ of subtype (E) can be expressed locally, using the convenient coordinates and convenient coframe, in the form (1.8), where $f_{2}=0, a_{2} \neq 0$ and $b_{2}=0$.

The proof is a modification of that of proposition 8.1 from [9] using the fact that $f A / f_{3}$ is a Riemannian invariant (cf. (3.14he)). Notice that the two cases in proposition 8.1 from [9] are reduced to one case only. In fact, if $f_{2}=a_{2}=0$, then, making substitution $\bar{y}=y+1$, we have $\bar{a}_{2} \neq 0$.

Now we study the "fine structure" of the partial differential equations

$$
\begin{equation*}
(A \alpha)_{y}^{\prime}+\beta_{x}^{\prime}=0 \tag{A1}
\end{equation*}
$$

and

$$
\begin{equation*}
R_{y}^{\prime}-\beta_{w}^{\prime}=0 \tag{A2}
\end{equation*}
$$

with $\beta \neq 0$, that is, we shall write (A1) and (A2) as a system of partial differential equations for functions of two variables only.

We substitute into (A1) the function $A \alpha$ in the form

$$
A \alpha=\frac{1}{2 f A}\left[\left(A^{2}\right)_{w}^{\prime}-2(A C)_{x}^{\prime}+\frac{A C}{A^{2}}\left(A^{2}\right)_{x}^{\prime}-H\left(A^{2}\right)_{y}^{\prime}\right]
$$

which follows from (1.12) and (1.10), and the function $\beta$ in the form $\beta=\lambda a_{0} / A^{2}$ (see (2.22)). Taking the common denominator $2 A^{4}(f A)^{2}$ and using (2.6he), (2.23he) and (2.28he), respectively, we obtain the nominator of the left-hand side of the equation (A1) as a linear combination of $c^{3}, c^{2} s, c^{2}, c s, c, s$ and 1 , where $c=\cosh (2 \lambda y)$ and $s=\sinh (2 \lambda y)$ in the hyperbolic case; $c=\cos (2 \lambda y)$ and $s=\sin (2 \lambda y)$ in the elliptic case. Each coefficient of this linear combination depends on $w$ and $x$ only, and thus it must vanish if (A1) is satisfied. This gives seven partial differential equations which are linear with respect to $a_{0 x}^{\prime}, a_{1 x}^{\prime}, a_{2 x}^{\prime}, a_{3 x}^{\prime}, V_{1}, V_{2}$ and $V_{3}$, where

$$
\left\{\begin{array}{l}
V_{1}=a_{1 w}^{\prime}-2 b_{1 x}^{\prime}-2 \lambda H a_{2},  \tag{5.3}\\
V_{2}=a_{2 w}^{\prime}-2 b_{2 x}^{\prime}+\epsilon 2 \lambda H a_{1}, \\
V_{3}=a_{3 w}^{\prime}-2 b_{3 x}^{\prime} .
\end{array}\right.
$$

Using the formula (2.24he) in the form

$$
\begin{equation*}
a_{0}^{2}=-\epsilon\left(a_{1}^{2}-a_{3}^{2}\right)-a_{2}^{2} \tag{5.4}
\end{equation*}
$$

and its derivative

$$
\begin{equation*}
a_{0} a_{0 x}^{\prime}=-\epsilon\left(a_{1} a_{1 x}^{\prime}-a_{3} a_{3 x}^{\prime}\right)-a_{2} a_{2 x}^{\prime}, \tag{5.5}
\end{equation*}
$$

we can eliminate the derivative $a_{0 x}^{\prime}$ in all equations. We obtain the final form of the equation (A1) as the system of partial differential equations

$$
\begin{equation*}
\sum_{i=1}^{3} a_{0} P_{\alpha}^{i} V_{i}+\sum_{i=1}^{3} Q_{\alpha}^{i} a_{i x}^{\prime}=0, \quad \alpha=1,2, \ldots, 7 \tag{5.6}
\end{equation*}
$$

where

$$
\begin{array}{lll}
P_{1}^{1}=2 a_{1} a_{2} f_{3}, & P_{1}^{2}=\left(a_{1}^{2}-\epsilon a_{2}^{2}\right) f_{3}, & P_{1}^{3}=-2 a_{1} a_{2} f_{1}, \\
P_{2}^{1}=\left(a_{1}^{2}-\epsilon a_{2}^{2}\right) f_{3}, P_{2}^{2}=-2 \epsilon a_{1} a_{2} f_{3}, & P_{2}^{3}=-\left(a_{1}^{2}-\epsilon a_{2}^{2}\right) f_{1}, \\
P_{3}^{1}=2 a_{2} a_{3} f_{3}, & P_{3}^{2}=\left(a_{1}^{2}-\epsilon a_{2}^{2}\right) f_{1}+2 a_{1} a_{3} f_{3}, & P_{3}^{3}=-2 a_{2} a_{3} f_{1}, \\
P_{4}^{1}=a_{1} a_{3} f_{3}, & P_{4}^{2}=-\epsilon a_{2}\left(a_{1} f_{1}+a_{3} f_{3}\right), & P_{4}^{3}=-a_{1} a_{3} f_{1}, \\
P_{5}^{1}=2 a_{1} a_{2} f_{3}, & P_{5}^{2}=-2 a_{1} a_{3} f_{1}-\left(\epsilon a_{2}^{2}+a_{3}{ }^{2}\right) f_{3}, P_{5}^{3}=-2 a_{1} a_{2} f_{1}, \\
P_{6}^{1}=\left(a_{2}^{2}+\epsilon a_{3}^{2}\right) f_{3}, P_{6}^{2}=-2 a_{2} a_{3} f_{1}, & P_{6}^{3}=-\left(a_{2}{ }^{2}+\epsilon a_{3}^{2}\right) f_{1}, \\
P_{7}^{1}=2 a_{2} a_{3} f_{3}, & P_{3}^{2}=-\left(\epsilon a_{2}^{2}+a_{3}^{2}\right) f_{1}, & P_{7}^{3}=-2 a_{2} a_{3} f_{1},
\end{array}
$$

$$
\begin{aligned}
& Q_{1}^{1}=-a_{0} a_{2} b_{3} f_{1}+a_{0} a_{2} b_{1} f_{3}+\left(a_{2}{ }^{2}-\epsilon a_{3}{ }^{2}\right) f_{1}{ }^{2}, \\
& Q_{2}^{1}=-a_{0}\left(a_{1} b_{3}-a_{3} b_{1}\right) f_{1}+a_{0} a_{1} b_{1} f_{3}+a_{1} a_{2} f_{1}{ }^{2}, \\
& Q_{3}^{1}=-a_{0} a_{2} b_{1} f_{1}-\epsilon a_{1} a_{3} f_{1}^{2}+2\left(a_{2}{ }^{2}-\epsilon a_{3}{ }^{2}\right) f_{1} f_{3}, \\
& Q_{4}^{1}=-a_{0} a_{3} b_{1} f_{3}+a_{1} a_{2} f_{1} f_{3}, \\
& Q_{5}^{1}=2 a_{0} a_{2} b_{1} f_{3}-\left(a_{2}{ }^{2}-\epsilon a_{3}{ }^{2}\right) f_{3}{ }^{2}+2 \epsilon a_{1} a_{3} f_{1} f_{3}, \\
& Q_{6}^{1}=\epsilon a_{0} a_{3} b_{3} f_{3}+\epsilon a_{1} a_{2} f_{3}{ }^{2}, \\
& Q_{7}^{1}=a_{0} a_{2} b_{3} f_{3}+\epsilon a_{1} a_{3} f_{3}{ }^{2} \text {, } \\
& Q_{1}^{2}=-a_{0}\left(a_{1} b_{3}-a_{3} b_{1}\right) f_{1}+a_{0} a_{1} b_{1} f_{3}-a_{1} a_{2} f_{1}{ }^{2}, \\
& Q_{2}^{2}=\epsilon a_{0} a_{2} b_{3} f_{1}-\epsilon a_{0} a_{2} b_{1} f_{3}-\left(a_{1}{ }^{2}-a_{3}{ }^{2}\right) f_{1}{ }^{2}, \\
& Q_{3}^{2}=a_{0} a_{1} b_{1} f_{1}+2 a_{0} a_{3} b_{1} f_{3}-a_{2} a_{3} f_{1}^{2}-2 a_{1} a_{2} f_{1} f_{3} \text {, } \\
& Q_{4}^{2}=-\left(a_{1}{ }^{2}-a_{3}{ }^{2}\right) f_{1} f_{3} \text {, } \\
& Q_{5}^{2}=-2 a_{0} a_{1} b_{3} f_{1}-a_{0} a_{3} b_{3} f_{3}+a_{1} a_{2} f_{3}{ }^{2}+2 \epsilon a_{2} a_{3} f_{1} f_{3}, \\
& Q_{6}^{2}=-a_{0} a_{2} b_{3} f_{1}+a_{0} a_{2} b_{1} f_{3}-\left(a_{1}{ }^{2}-a_{3}{ }^{2}\right) f_{3}{ }^{2} \text {, } \\
& Q_{7}^{2}=-a_{0} a_{3} b_{3} f_{1}-a_{0}\left(a_{1} b_{3}-a_{3} b_{1}\right) f_{3}+\epsilon a_{2} a_{3} f_{3}{ }^{2}, \\
& Q_{1}^{3}=-a_{0} a_{2} b_{1} f_{1}+\epsilon a_{1} a_{3} f_{1}{ }^{2}, \\
& Q_{2}^{3}=-a_{0} a_{1} b_{1} f_{1}-a_{2} a_{3} f_{1}{ }^{2}, \\
& Q_{3}^{3}=-2 a_{0} a_{2} b_{3} f_{1}+\left(\epsilon a_{1}^{2}-a_{3}^{2}\right) f_{1}^{2}+2 \epsilon a_{1} a_{3} f_{1} f_{3}, \\
& Q_{4}^{3}=-a_{0} a_{1} b_{3} f_{1}-a_{2} a_{3} f_{1} f_{3}, \\
& Q_{5}^{3}=a_{0} a_{2} b_{3} f_{3}-\epsilon a_{1} a_{3} f_{3}^{2}+2\left(\epsilon a_{1}^{2}+a_{2}^{2}\right) f_{1} f_{3} \text {, } \\
& Q_{6}^{3}=-a_{0} a_{3} b_{3} f_{1}-\epsilon a_{0}\left(a_{1} b_{3}-a_{3} b_{1}\right) f_{3}-\epsilon a_{2} a_{3} f_{1} f_{3}{ }^{2}, \\
& Q_{7}^{3}=-a_{0} a_{2} b_{3} f_{1}+a_{0} a_{2} b_{1} f_{3}-\left(\epsilon a_{1}{ }^{2}+a_{2}{ }^{2}\right) f_{3}{ }^{2} .
\end{aligned}
$$

Next, we substitute into (A2) the function $R$ in the form

$$
R=\frac{1}{2 f A}\left[\left(f^{2}+C^{2}\right)_{x}^{\prime}+H\left(h+(A C)_{y}^{\prime}\right)-\frac{A C}{A^{2}}\left(A^{2}\right)_{w}^{\prime}\right],
$$

which we obtain from $(1.13)_{1}$, (1.12) and (1.10), and the function $\beta$ in the form $\beta=\lambda a_{0} / A^{2}$. By the same argument as that for the previous equation (A1), we obtain once more seven partial differential equations.

They are now linear with respect to $a_{0 w}^{\prime}, a_{1 w}^{\prime}, a_{2 w}^{\prime}, a_{3 w}^{\prime}, W_{1}, W_{2}$ and $W_{3}$, where

$$
\left\{\begin{align*}
W_{1} & =-\epsilon \frac{1}{\lambda} \varphi_{1 x}^{\prime}+2 \lambda H b_{2}  \tag{5.7}\\
W_{2} & =\frac{1}{\lambda} \varphi_{2 x}^{\prime}+2 \epsilon \lambda H b_{1} \\
W_{3} & =\frac{1}{\lambda} \varphi_{3 x}^{\prime}+H h
\end{align*}\right.
$$

Using (5.4) and the formula for $a_{0 w}^{\prime}$ similar to (5.5), we can also eliminate the derivative $a_{0 w}^{\prime}$ in all equations. We obtain the final form of the equation (A2) as a system of partial differential equations analogous to (5.6):

$$
\begin{equation*}
\sum_{i=1}^{3} a_{0} P_{\alpha}^{i} W_{i}-\sum_{i=1}^{3} Q_{\alpha}^{i} a_{i w}^{\prime}=0, \quad \alpha=1,2, \ldots, 7 \tag{5.8}
\end{equation*}
$$

The following proposition will be crucial for reducing our partial differential equations to essentially independent ones.

Proposition 5.2. The rank of the matrix $\left[P_{\alpha}^{i}, Q_{\alpha}^{i}\right]$ is at most two.

Proof. Since $a_{2} \neq 0$ and $b_{2}=f_{2}=0$, we have from (2.51)

$$
\begin{equation*}
\varphi_{4}=\epsilon \frac{\lambda\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{2}} \tag{5.9}
\end{equation*}
$$

and hence we have

$$
\left\{\begin{align*}
b_{1} & =\frac{a_{2}^{2} f_{3}+\epsilon a_{1}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{0} a_{2}}  \tag{5.10}\\
b_{3} & =\frac{a_{2}^{2} f_{1}+\epsilon a_{3}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{0} a_{2}}
\end{align*}\right.
$$

Substituting from (5.10) for $b_{1}$ and $b_{3}$ in the entries of the matrix $\left[Q_{\alpha}^{i}\right]$,
we see that

$$
\begin{aligned}
& {\left[P_{\alpha}^{3}\right]=-\frac{f_{1}}{f_{3}}\left[P_{\alpha}^{1}\right]} \\
& {\left[Q_{\alpha}^{1}\right]=\epsilon \frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}}\left[P_{\alpha}^{1}\right]-\epsilon f_{3}\left[P_{\alpha}^{2}\right]} \\
& {\left[Q_{\alpha}^{2}\right]=-\frac{f_{1}^{2}-f_{3}^{2}}{f_{3}}\left[P_{\alpha}^{1}\right]+\epsilon \frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}}\left[P_{\alpha}^{2}\right]} \\
& {\left[Q_{\alpha}^{3}\right]=-\epsilon \frac{f_{1}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{2} f_{3}}\left[P_{\alpha}^{1}\right]+\epsilon f_{1}\left[P_{\alpha}^{2}\right]}
\end{aligned}
$$

which prove the assertion.
Corollary 5.3. Each system of partial differential equations (5.6) or (5.8) contains at most two linearly independent equations.

Thus, the equations (A1) and (A2) are essentially reduced to four partial differential equations in two variables. We shall see later that, as in [9], we can make an additional reduction to only two equations (one of the form (5.6) and one of the form (5.8)).

Proposition 5.4. The following algebraic formulas are consequences of the algebraic equations from theorem 2.8 and of the assumptions of proposition 5.1:

$$
\begin{equation*}
\varphi_{1}=\nu a_{1}, \quad \varphi_{2}=\epsilon \nu a_{2}, \quad \varphi_{3}=-\epsilon \nu a_{3} \tag{5.11}
\end{equation*}
$$

where

$$
\begin{equation*}
\nu=\frac{\lambda\left[a_{2}^{2}\left(f_{1}^{2}-f_{3}^{2}\right)-\epsilon\left(a_{1} f_{3}-a_{3} f_{1}\right)^{2}\right]}{a_{0}^{2} a_{2}^{2}}, \quad \nu=\epsilon \frac{\varphi_{0}}{a_{0}}, \tag{5.12}
\end{equation*}
$$

$$
\begin{equation*}
a_{0}^{2}=-\epsilon\left(a_{1}^{2}-a_{3}^{2}\right)-a_{2}^{2} . \tag{5.13}
\end{equation*}
$$

Further, $f_{2}=0$ and

$$
\left\{\begin{array}{l}
b_{1}=\frac{a_{2}^{2} f_{3}+\epsilon a_{1}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{0} a_{2}}  \tag{5.14}\\
b_{2}=0 \\
b_{3}=\frac{a_{2}^{2} f_{1}+\epsilon a_{3}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{0} a_{2}}
\end{array}\right.
$$

(5.15) $\quad h=-\epsilon \frac{2 \lambda\left(a_{1} f_{1}-a_{3} f_{3}\right)}{a_{0}}, \quad \varphi_{4}=\epsilon \frac{\lambda\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{2}}, \quad \varphi_{5}=2 \varphi_{4}$.

Conversely, if $a_{1}, a_{2}, a_{3}, f_{1}$ and $f_{3}$ are arbitrary functions, and if the other basic functions are defined as above, then all algebraic equations of theorem 2.8 hold.

Proof. We show only the necessity of (5.11)-(5.15). The sufficiency will be proved by the direct check. The equations $(2.44)_{3}$ and $(2.44)_{5}$ imply $a_{1} \varphi_{2}-\epsilon a_{2} \varphi_{1}=0$ and $a_{2} \varphi_{3}+a_{3} \varphi_{2}=0$. Hence the formulas (5.11) hold with some function $\nu=\nu(w, x)$ of the variables $w$ and $x$. Substituting (5.11) and $(5.10)_{1}$ into $(2.44)_{1}$, and using (2.24he), we obtain $(5.12)_{1}$. The formula (5.13) is a direct consequence of (2.24he). The formulas $(5.14)_{1,3}$ and $(5.15)_{2}$ follow from $b_{2}=f_{2}=0$ as shown in the proof of proposition 5.2. Next, from (5.11), (2.5he) and (2.24he), we have $\varphi_{0}{ }^{2}=\nu^{2} a_{0}{ }^{2}$. Here, the relation (4.3) implies that $\epsilon\left(f_{1}{ }^{2}-f_{3}{ }^{2}\right)$ is negative because the discriminant $\Delta^{\prime}$ is negative, hence $\epsilon \nu$ is negative. On the other hand, (4.4) together with (4.3) implies that $a_{0} \varphi_{0}$ is negative. Hence we obtain $(5.12)_{2}$. We obtain $(5.15)_{1}$ from $(2.47)$ and $f_{2}=0$. Finally, $(5.15)_{3}$ is the same as (2.48).

We need later the relation

$$
\begin{equation*}
\nu=\frac{\lambda\left(f_{1} b_{3}-f_{3} b_{1}\right)}{a_{0} a_{1}} \tag{5.16}
\end{equation*}
$$

which follows from (5.14) and (5.12).
Now let us return to the system of partial differential equations (5.6) and (5.8). Specifying corollary 5.3 , we see easily that the system (5.6) reduces to two partial differential equations

$$
\begin{equation*}
a_{0} V_{2}-\epsilon f_{3} a_{1 x}^{\prime}+\epsilon \frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}} a_{2 x}^{\prime}-f_{1} a_{3 x}^{\prime}=0 \tag{5.17}
\end{equation*}
$$

and

$$
\begin{align*}
a_{0} f_{3} V_{1} & +\frac{a_{0}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{2}} V_{2}-a_{0} f_{1} V_{3}+ \\
& +\left[\epsilon\left(\frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}}\right)^{2}-\left(f_{1}^{2}-f_{3}^{2}\right)\right] a_{2 x}^{\prime}=0 \tag{5.18}
\end{align*}
$$

The system (5.8) reduces to two analogous equations

$$
\begin{equation*}
a_{0} W_{2}+\epsilon f_{3} a_{1 w}^{\prime}-\epsilon \frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}} a_{2 w}^{\prime}+f_{1} a_{3 w}^{\prime}=0 \tag{5.19}
\end{equation*}
$$

and

$$
\begin{align*}
a_{0} f_{3} W_{1} & +\frac{a_{0}\left(a_{1} f_{3}-a_{3} f_{1}\right)}{a_{2}} W_{2}-a_{0} f_{1} W_{3}+ \\
& -\left[\epsilon\left(\frac{a_{1} f_{3}-a_{3} f_{1}}{a_{2}}\right)^{2}-\left(f_{1}^{2}-f_{3}^{2}\right)\right] a_{2 w}^{\prime}=0 . \tag{5.20}
\end{align*}
$$

Using (5.3), (5.7), (5.11), (5.14) and (5.16), we see, after lengthy but routine calculations, that (5.18) and (5.20) are consequences of (5.17) and (5.19).

Substituting $(5.3)_{2}$ and $(5.7)_{2}$ into (5.17) and (5.19), respectively, and using $(5.11)_{2}$, we have

$$
\left\{\begin{array}{l}
a_{0} a_{2 w}^{\prime}+\epsilon 2 \lambda H a_{0} a_{1}-\epsilon a_{2} f_{3}\left(\frac{a_{1}}{a_{2}}\right)_{x}^{\prime}+\epsilon a_{2} f_{1}\left(\frac{a_{3}}{a_{2}}\right)_{x}^{\prime}=0  \tag{5.21}\\
a_{0}\left(\nu a_{2}\right)_{x}^{\prime}-2 \lambda^{2} H a_{0} b_{1}+\lambda a_{2} f_{3}\left(\frac{a_{1}}{a_{2}}\right)_{w}^{\prime}-\lambda a_{2} f_{1}\left(\frac{a_{3}}{a_{2}}\right)_{w}^{\prime}=0
\end{array}\right.
$$

Further, due to $(5.15)_{1}$, we have the relation

$$
\begin{equation*}
2 \lambda\left(a_{1} f_{1}-a_{3} f_{3}\right)=-\epsilon a_{0} H_{x}^{\prime} \tag{5.22}
\end{equation*}
$$

Introducing new functions $u=u(w, x)$ and $v=v(w, x)$ of the variables $w$ and $x$ such that

$$
\begin{equation*}
a_{1}=u a_{2}, \quad a_{3}=v a_{2}, \quad-\epsilon\left(u^{2}-v^{2}\right)>0 \tag{5.23}
\end{equation*}
$$

we rewrite (5.21) in the form

$$
\left\{\begin{array}{l}
a_{0} a_{2 w}^{\prime}+\epsilon 2 \lambda H a_{0} a_{1}-\epsilon a_{2} f_{3} u_{x}^{\prime}+\epsilon a_{2} f_{1} v_{x}^{\prime}=0  \tag{5.24}\\
a_{0}\left(\nu a_{2}\right)_{x}^{\prime}-2 \lambda^{2} H a_{0} b_{1}+\lambda a_{2} f_{3} u_{w}^{\prime}-\lambda a_{2} f_{1} v_{w}^{\prime}=0
\end{array}\right.
$$

Here, from (5.12)-(5.14), we get

$$
\left\{\begin{array}{l}
a_{0}=\sqrt{-\epsilon\left(u^{2}-v^{2}\right)-1} a_{2}  \tag{5.25}\\
b_{1}=\frac{f_{3}+\epsilon u\left(u f_{3}-v f_{1}\right)}{\sqrt{u^{2}-v^{2}-1}} \\
\nu=\frac{\lambda\left[f_{1}^{2}-{\left.f_{3}^{2}-\epsilon\left(u f_{3}-v f_{1}\right)^{2}\right]}_{\left(u^{2}-v^{2}-1\right) a_{2}^{2}}\right.}{}=\frac{}{2}
\end{array}\right.
$$

where we normalize the signs of $a_{2}$ and $a_{0}$ to make them positive.
Let now $u, v$ and $H$ be arbitrary analytic functions. Substituting for $a_{0}$ from $(5.25)_{1}$ into (5.22) and into $(5.24)_{1}$, and solving them with respect to $f_{1}$ and $f_{3}$, we can express $f_{1}$ and $f_{3}$ in the form

$$
\left\{\begin{array}{l}
f_{1}=g_{1} a_{2 w}^{\prime}+g_{2} a_{2}+g_{3}  \tag{5.26}\\
f_{3}=h_{1} a_{2 w}^{\prime}+h_{2} a_{2}+h_{3}
\end{array}\right.
$$

where $g_{i}$ 's and $h_{i}$ 's are known functions. Substituting (5.26) into (5.24) ${ }_{2}$ which has been transformed by (5.25), we obtain a partial differential equation of the form

$$
\begin{equation*}
a_{2 w x}^{\prime \prime}=\Psi\left(a_{2 w}^{\prime}, a_{2 x}^{\prime}, a_{2}, w, x\right) \tag{5.27}
\end{equation*}
$$

where $\Psi$ is a fixed analytic function of five variables. The general solution of (5.27) depends on two arbitrary (analytic) functions of one variable. Thus, the generic family of metrics of subtype (E) depends on three arbitrary functions of two variables, namely, $u, v$ and $H$.

Now, we can go further and prove that even the local isometry classes of our metrics still depend essentially on three functions. The proof is a modification of that of theorem 8.5 from [9]. We use the fact that $f A / f_{3}$ is a Riemannian invariant (see (3.14he)) and that the hyperbolic cosine function and the cosine function are even functions.

THEOREM 5.5. The local isometry classes of metrics of subtype (E) are parametrized by three arbitrary functions of two variables modulo two arbitrary functions of one variable.

The equation (5.27) can not be solved explicitly, in general. Yet, we give here an explicit family of the metrics of subtype (E).

Example 5.6. Consider the "singular" case $a_{2}=0$ of proposition 5.4. Then we have

$$
\begin{cases}\varphi_{1}=\nu a_{1}, & \varphi_{2}=a_{2}=0, \quad \varphi_{3}=-\epsilon \nu a_{3}  \tag{5.28}\\ \varphi_{0}=\epsilon \nu a_{0}, & \varphi_{5}=2 \varphi_{4}\end{cases}
$$

and

$$
\begin{equation*}
b_{2}=f_{2}=0 \tag{5.29}
\end{equation*}
$$

From $(2.45)_{2}$ we see that there is a function $\xi=\xi(w, x)$ of the variables $w$ and $x$ such that

$$
\begin{equation*}
f_{1}=\xi a_{1}, \quad f_{3}=\xi a_{3} \tag{5.30}
\end{equation*}
$$

Hence, using (5.22) and (2.51), we have

$$
\begin{align*}
a_{0} h & =-\epsilon 2 \lambda \xi\left(a_{1}^{2}-a_{3}^{2}\right),  \tag{5.31}\\
b_{1} & =\frac{a_{1} \varphi_{4}}{\lambda a_{0}}, \quad b_{3}=\frac{a_{3} \varphi_{4}}{\lambda a_{0}} .
\end{align*}
$$

Finally, we have

$$
\begin{equation*}
a_{0}^{2}=-\epsilon\left(a_{1}^{2}-a_{3}^{2}\right), \tag{5.33}
\end{equation*}
$$

and, from $(2.43)_{1}$ or $(2.44)_{1}$, we deduce

$$
\begin{equation*}
\frac{\varphi_{4}^{2}}{\lambda a_{0}^{2}}+\lambda \xi^{2}=-\epsilon \nu \tag{5.34}
\end{equation*}
$$

Here $a_{1}, a_{3}, \xi$ and $\varphi_{4}$ are arbitrary functions of the variables $w$ and $x$. Conversely, if $a_{1}, a_{3}, \xi$ and $\varphi_{4}$ are arbitrary functions of the variables $w$ and $x$, and if the other basic functions are given by (5.28)-(5.34), then all algebraic equations mentioned in theorem 2.8 are satisfied.

In addition, from (5.31) and (5.33), we get

$$
\begin{equation*}
h=2 \lambda \xi a_{0}, \quad h=H_{x}^{\prime} \tag{5.35}
\end{equation*}
$$

Further, a careful check shows that the system of partial differential equations (5.6) and (5.8) can be now reduced, instead of the form (5.21), to the form

$$
\begin{gather*}
a_{0} V_{2}-\epsilon\left(f_{3} a_{1 x}^{\prime}-f_{1} a_{3 x}^{\prime}\right)=0  \tag{5.36}\\
a_{0} W_{2}+\epsilon\left(f_{3} a_{1 w}^{\prime}-f_{1} a_{3 w}^{\prime}\right)=0 \tag{5.37}
\end{gather*}
$$

All other partial differential equations are consequences of (5.36) and (5.37). Putting $U=a_{3} / a_{1}$, we can rewrite (5.36) and (5.37) in the form

$$
\begin{align*}
& 2 \lambda H a_{0}+\xi a_{1} U_{x}^{\prime}=0  \tag{5.38}\\
& 2 H \varphi_{4}+\xi a_{1} U_{w}^{\prime}=0
\end{align*}
$$

Then we have the following explicit family of solutions satisfying the equations (5.38) and (5.39) and the condition (5.35). Choose $U$ and $H$ as arbitrary functions of the variables $w$ and $x$, and put
(5.40) $\begin{cases}a_{1}=-\epsilon \frac{h U_{x}^{\prime}}{4 \lambda^{2} H\left(U^{2}-1\right)}, & a_{3}=a_{1} U, \quad a_{0}=a_{1} \sqrt{\epsilon\left(U^{2}-1\right)}, \\ \xi=-\frac{2 \lambda H \sqrt{\epsilon\left(U^{2}-1\right)}}{U_{x}^{\prime}}, & \varphi_{4}=-\frac{h U_{w}^{\prime}}{4 \lambda H \sqrt{\epsilon\left(U^{2}-1\right)}}, \quad h=H_{x}^{\prime} .\end{cases}$

Here we always assume $U_{x}^{\prime} \neq 0$ and $\epsilon\left(U^{2}-1\right)>0$. (Also, we normalize the signs of $a_{1}, a_{3}$ and $a_{0}$ to make them all positive.) Then the function $\nu$ is calculated from (5.34) and remaining coefficients are given by (5.28)(5.30) and (5.32). This defines the wanted class of metrics.

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